





- Transition Edge Sensors with FDM based on SQUIDs is baseline for SAFARI
- Active European Consortium Working on Ultra-Low-Noise TES Arrays since 2009
- Current principal members are Cambridge Univ., Cardiff Univ. and SRON
- Supported by ESA, largely in context of SAFARI, but not SAFARI specific
- No SAFARI-specific UK funding during this time, but programme has drawn heavily on large STFC investment in infrastructure in superconducting detector technology

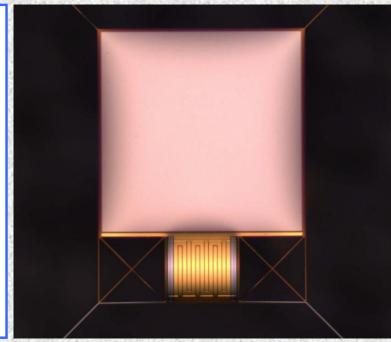


- In 2013 status changed from TRP to CTP, with UK and SRON developments running in parallel, rather than a single funded programme
- Very strong collaboration continues with SRON on all aspects of SAFARI instrument
- Entering last phase of funded work on CTP, which will finish Dec 2016
- Over last 10 years numerous workshops, meetings and progress reviews
- UK has substantial capability in low-noise FIR TES arrays (NEP 10⁻¹⁷ to 10⁻¹⁹ WHz^{-1/2})

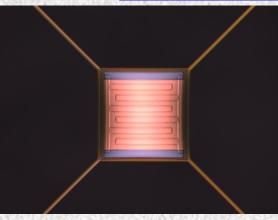
Transition Edge Sensors:

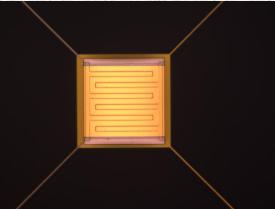
 β -phase Ta absorber with Tc of 860 mK 200 nm SiN (high heat capacity due to TLSs) MoAu bilayers with Tc of 110 mK $G \sim 0.2 \text{ pW K}^{-1}$ NEP $\sim 2 \text{ to } 4 \text{ x } 10^{-19} \text{ WHz}^{-1/2}$ $\tau \sim 10 \text{ mS}$ Psat $\sim 20 \text{ fW}$

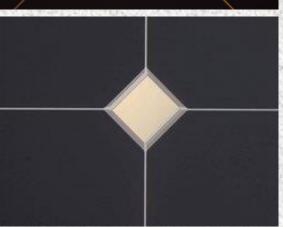
Low Psat and NEP means measurement challenges



Bilayer technologies needed to achieve Tc - all make good TESs







Cambridge MoCu – 200nm (40/106nm)

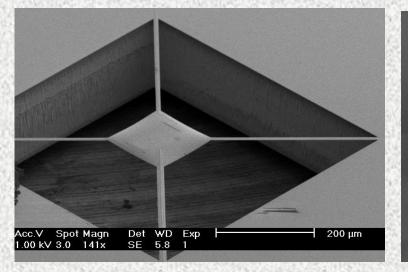
- Especially difficult to manufacture
- Need SiO₂ passivation layer to protect Cu
- High stress thin (200nm) nitride $<1\mu$ m curvature

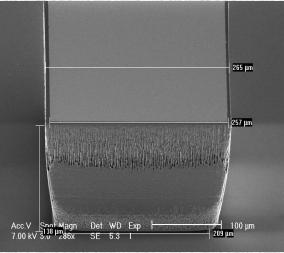
Cambridge MoAu – 200nm

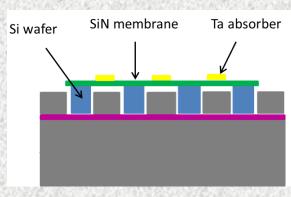
- Self passivating low C
- Good inter-diffusion stability
- Low stress good for thin nitride < 100nm curvature

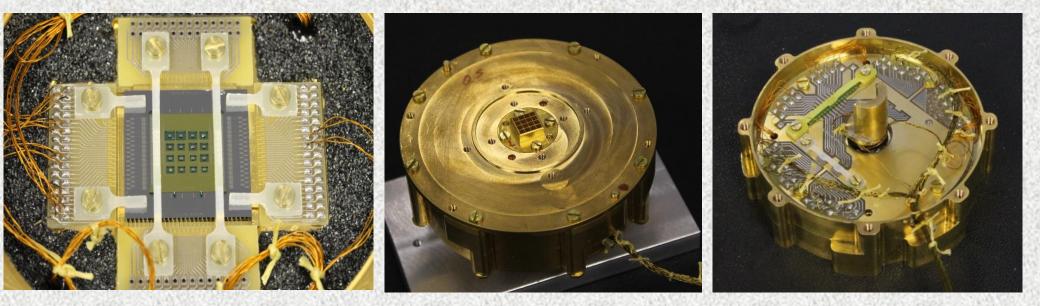
SRON TiAu - 500nm (16/85)

- Self passivating
- Good inter-diffusion stability







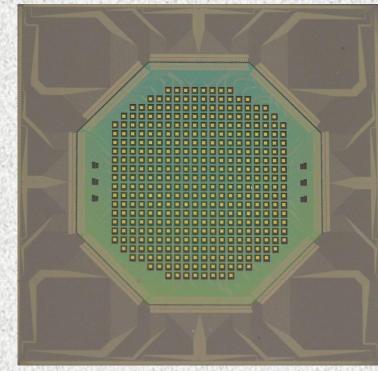


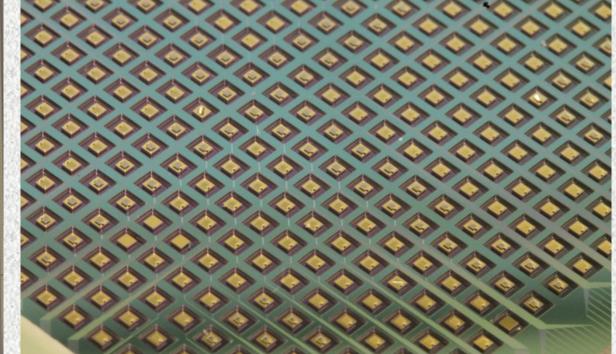
Small-format 16 element test arrays with integrated optical backshorts Few-mode, 10-20, lightpipe coupling (developed frabrication technology for S and L Band)

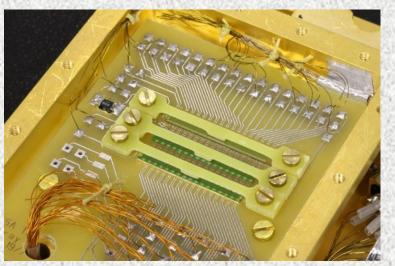
First phase (TRP):

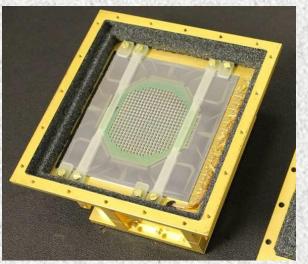
- UK work concentrated on large-format L-Band imaging arrays (200-100 μm)
- SRON concentrated on large-format S-Band imaging arrays (30-60 μm)
- No M-Band work (60-110 μ m) most difficult
- Development of fabrication techniques
- Exploration of basic device physics (what determines noise, speed, etc.)
- Demonstration of laboratory-based measurement techniques (variable temperature load, passband defining optical filters)
- Only limited optical tests proven, but very difficult
- Few-mode system and so work was needed to set up data analysis models of test system

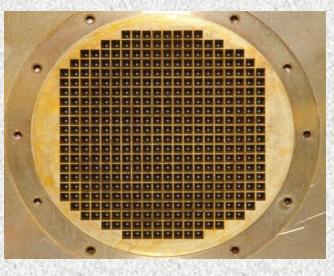
ESA TRP L Band 388 pixel imaging array as off 2013:









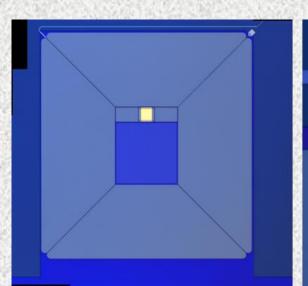


Second phase (CTP):

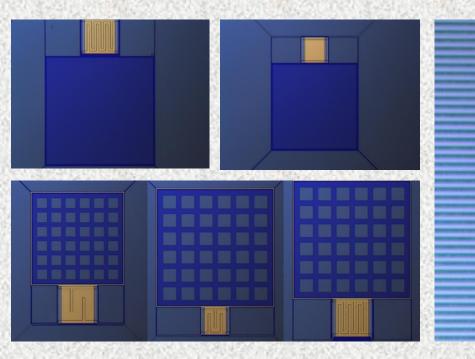
- UK worked on small-format L-Band, M Band and S-Band arrays 200-30 μm
- Extensive work on modelling, device processing, more sophisticated designs
- Development of profiled lightpipe manufacture, camera integration, metrology
- Measurements of dark noise and device dynamics
- Accurate optical efficiency measurements development of bandpass optical filters
- Designed and fabricated complete sets of `high G' devices having same physical forms of `low-G' devices.
- Now have spectral and spatial measurements using room temperature FTS and beam scanning system (test systems set up at Cardiff)
- Excellent imaging array technology available for utilitarian purposes as by-product

TES array chips:

- Extreme range of leg geometries: 1500µm long and 1.5µm wide 200 µm long and 100µm wide 4um long and 500 nm wide
- Bilayer variations (size, Au bars)
- Absorber variations (meshed)
- Au rim / no rim





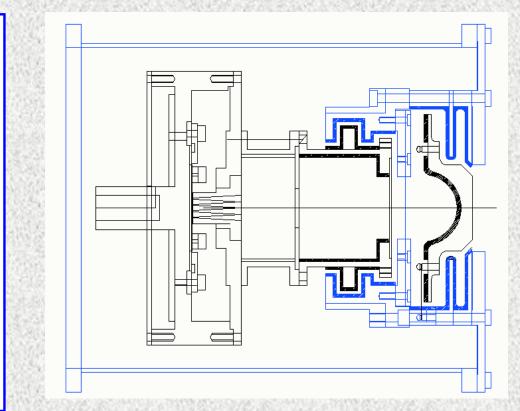


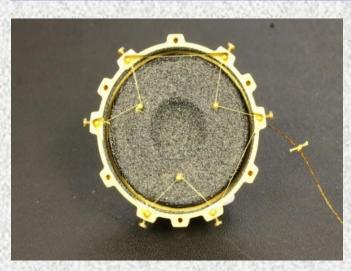
Microstrip Nb wiring:

- Fully integrated process Nb/SiO₂/Nb track width 2µm space between tracks 2µm 250 tracks/mm density
- Breakout to standard wiring on legs
- Excellent alignment on legs

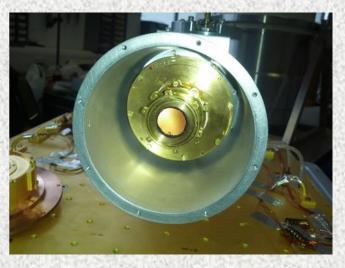
Optical efficiency measurements(Cambridge)

- Thermal and bandpass optical filters
- Aperture defined throughput
- Detailed modal analysis
- Clean few-mode illumination
- Black absorber throughout
- Variable temperature hot load: 3.5 – 25 K range (3 heaters) 200 µK temperature stability
 < 10 s time constant set by wire
- Extremely high straylight rejection



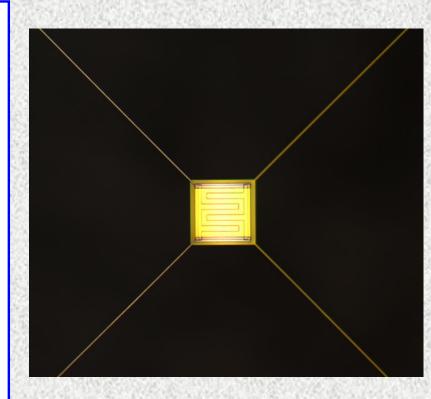


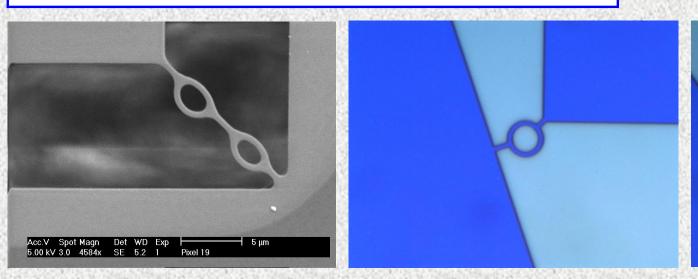




Many development opportunities available.

- Improved sensitivity lower NEPs, uniformity
- Move to Sol rather than SiN to avoid heat capacity of TLSs and increase speed, higher uniformity
- Ballistic phononic legs (optical lithography and EBL)
- Excellent results already demonstrated for quantum thermal conductance limited TESs





Coming year (2016):

- Manufacture and test demonstration of M Band linear arrays (60-110 μm)
- Concentrate on mass-producible linear arrays for grating readout
- Refine profiled lightpipe manufacturing techniques and camera assembly techniques
- Understand optical efficiency measurements in greater detail

Proposed that UK should contribute to SAFARI mission:

- UK expertise in TES technology can make a major contribution to the SAFARI instrument (not just device processing, which is only a small part of an imaging spectroscopic sub-system, also need metrology and array integration)
- Current working arrangement that UK would provide detector arrays (not just devices) for at least one, possibly two, wavebands: M and/or S Band being the most likely
- Excellent if the UK could make a key high-tech contribution to a major, high profile space mission, working with international colleagues
- SRON and potential US partners have agreed to work with us in context of the UK providing detector technology for SAFARI
- Would place UK in position of being a major player in developing and providing technology for the mission and a beneficiary of the resulting science